SwissCube Project Phase D Qualification and Flight Acceptance Review, March 9, 2009

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### **Science Objectives**

Measure the airglow emission in the upper atmosphere at 100 km altitude to :

- Demonstrate the feasibility of using airglow as a basis for a low-cost earth sensor
- Validate the established airglow model or bring additional information about airglow dependence on
  - → latitude
  - $\rightarrow$  altitude
  - $\rightarrow$  local solar time



nightglow and aurora borealis









#### Optical Alignment Test

#### Payload Design

CMOS detector MT9V032 188 x 120 pixels, pixel size = 24  $\mu$ m resolution = 0.16°/pixel FOV = 18.8 x 25°

#### focusing optics triplet design with OTS components

support structure titanium

#### baffle

solar exclusion angle = 30° attenuation factor = 10<sup>-4</sup> vanes: stainless steel spacers: aluminium

> closing cap aluminium

filter CWL 767 nm FWHM 20 nm

#### payload board

microcontroller MSP430F1611 CMOS detector MT9V032 temperature sensor LM94022 oscillator HC-49/US SMD RAM R1LV0416CSB-7LI

- Purpose: test robustness of the opto-mechanical system of the payload against vibrations, shocks and temperature variations
- Tested Requirement:

#### Field of view

The payload may have a FOV of at least 20°.

#### **Angular resolution**

The payload may measure airglow with an angular resolution of at least 0.3°.

 $\rightarrow$  Maximum spot diameter = 48 µm (2 pixels)



- Test Setup:
  - Test bench for payload alignment test
- Limitations of the PL test bench
  - intensity of the laser spot
  - angle of incidence of the laser
- Updated success criteria

 $\rightarrow$  spot size and intensity distribution have to remain constant

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- Test results for the QM:
  - variation of the spot diameter  $\leq$  pixel
  - variation of the intensity distribution  $\leq$  pixel

Angle of incidence	-8.5 [deg]	0 [deg]	8.5 [deg]
Full image after optical calibration		•	
Spot after optical calibration		-	÷
Full image after QM test campaign	-	•	
Spot after QM test campaign	0.2000		

- Test results for the FM:
  - variation of the spot diameter  $\leq$  pixel
  - variation of the intensity distribution  $\leq$  pixel

Angle of incidence	-8.5 [deg]	0 [deg]	8.5 [deg]
Full image after optical calibration			
Spot after optical calibration	-		
Full image after QM test campaign	•	-	
Spot after QM test campaign			

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First image after power up is grey (noise)

**Optical Alignment Test** 

- After each power up of the payload, two pictures will be taken and only the second one will be downloaded.
- Analysis of the spot for PL alignment tests should be done more accurately
  - Intensity distribution of the spot should be done to determine the RMS spot diameter
- Characterisation of the CMOS detector has been done on a DM, but should be repeated for the FM prior to the launch

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# **Conclusions and Future Work**

- Payload QM and FM successfully passed their test campaign
- Additional work related to the payload and science mission :
  - programming of a tool to analyse the images (ongoing)
  - programming of a tool to determine when an image has to be taken





- Purpose: characterise the dark signal of the CMOS detector vs. temperature
- Tested Requirement:

#### Signal-to-Noise-Ratio

**Optical Alignment Test** 

The payload shall allow taking science measurements with a SNR of at least 3 for limb measurements.

 $\rightarrow$  Maximum mean dark signal = 350 kHz

- Test Setup:
  - Thermal Chamber
  - Dark room

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### Test: Performance of the CMOS Detector vs. Temperature

- Test condition:
  - -15°C to °70°C
  - 2 different register configurations of the detector

#### airglow signal

at night: 8k - 40k e<sup>-</sup>/s at day: 470k - 4M e<sup>-</sup>/s

- Test results:
  - mean DS < min airglow signal if T < 40°C,</li>
  - PL board is operational between -15° and 70°C





### **Aiglow Model**

#### Intensity of the airglow Integration time EarthModel\_FV - Emission type-----Intensity Scale DCR Integration Time Saturation Intensity (photons/pixel) Mean Dark Count Rate (Hz) Integration time (s) Mean Emission Min Emission 4 • 4 Max Emission Simulate SPADs \* 4 4 + 4 Þ. Simulate Aurora 4 • Time of Observation Simulation of DS Simulation of the 01:00 Moon and aurora 02:00 03:00 04:00 05:00 Intensity of 06:00 Image of the 07:00 the airglow 08:00 airglow as seen 09:00 Local solar time from the satellite, 0 10:00 -0 11:00 here from a GEO 12:00 0 13:00 0 14:00 0 15:00 16:00 17:00 0 18:00 0 19:00 0 20:00 0 21:00 0 22:00 0 23:00 24:00 500 Image 1 Save Image

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- Payload may be a technology demonstrator of an earth sensor based on airglow
  - Observes the emission at 762 nm with a bandwidth between 10 nm and 40 nm
  - Has a spatial resolution of at least 0.3° and a FOV of 20°
  - Can perform science mission with the sun no closer than 30° from its boresight.
- Physical and electrical constraints

**Optical Alignment Test** 

- Volume:  $30 \times 30 \times 65 \text{ mm}^3$  for the optics  $80 \times 35 \times 15 \text{ mm}^3$  for the payload board
- Mass: < 50 g
- Peak Power: < 450 mW during 30 s for each image</li>
- Additional design driver
  - The PL board is not a critical element of the SwissCube satellite

 $\rightarrow$  no redundancy has been taken into account for this subsystem



optics of the AIRES earth sensor

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During the first 3 month one image each 4.5 days:

**Optical Alignment Test** 

- 5 images of dayglow/nightglow measured at limb/nadir
- Total: 20 measurements, cycle repetition of 18 days
- After 3 month: observation of variation of emission intensity depending on latitude
  - Dayglow/nightglow above 85° N/S
  - Dayglow/nightglow between 40° and 50° N/S
  - Dayglow/nightglow between 5° N and 5° S
  - Total : 10 measurements, cycle of repetition of 45 days
    8 measurements per latitude in one year

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### **Operational Scenario: PL Board**

- PL board always turned on for housekeeping
- Detector turned on only when science observations are carried out
- Science observations are triggered by EPS
- Power consumption:
  - 8 mW when no science observations are performed
  - < 450 mW during science observations</li>



### **Design Description: Optical System**

- Triplet design with off-the-shelf components
- FOV 18.8° x 25°

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- Resolution 0.16°/pixel
- Baffle for a solar exclusion angle of 30° with an attenuation factor of 10<sup>-4</sup>
- Filter with a central wavelength at 767 nm and a bandwidth of 20 nm







- Microcontroller MSP430F1611
  - Operate the detector
  - Communicate with the EPS

**Optical Alignment Test** 

- Read temperature sensors
- CMOS Detector MT9V032
  - Capture images of the airglow
- Temperature Sensor LM94022
  - Used for dark signal correction
- Oscillator HC-49/US SMD
  - Provide clock reference for the CMOS detector
- RAM R1LV0416CSB-7LI
  - Store images until transmission to ground station





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# Design Description: Expected Airglow Images

#### **Limb Observations**



#### Zenith Observations

